Excitation spectra from angular momentum projection of Hartree-Fock states and the configuration-interaction shell-model

Joshua T. Staker and Calvin W. Johnson Department of Physics, San Diego State University, 5500 Campanile Drive, San Diego, CA 92182-1233

We make numerical comparison of spectra from angular-momentum projection on Hartree-Fock states with spectra from configuration-interaction nuclear shell-model calculations, all carried out in the same model spaces (in this case the sd, lower pf, and p -sd_{5/2} shells) and using the same input Hamiltonians. We find, unsurprisingly, that the low-lying excitation spectra for rotational nuclides are well reproduced, but the spectra for vibrational nuclides, and more generally the complex specta for odd-A and odd-odd nuclides are less well reproduced in detail.

I. INTRODUCTION

In Hartree-Fock (HF) calculations, one approximates the ground state by a single Slater determinant (antisymmeterized product of single-particle wavefunctions) and minimizes the energy, reducing the many-body problem to an effective one-body problem [\[1\]](#page-11-0). The Hartree-Fock solution can break exact symmetries, such as rotational invariance, and indeed is often more effective if it breaks exact symmetries. Restoring broken symmetries adds correlations that further lower the energy and is a useful approach for many problems.

In this paper we project states of good angular momentum from a Hartree-Fock state, and compare the resulting ground state energies and, especially, excitation spectra to exact results from equivalent full configuration-interaction diagonalization in a shell-model basis. Such comparisons help us to understand the accuracy and limitations of the projected Hartree-Fock approximation (PHF), as well as potentially providing a useful shortcut to nuclear structure.

PHF is not new [\[2\]](#page-11-1) and can be found in a number of variants. Most of these are inspired by the Nilsson model [\[3,](#page-11-2) [4](#page-11-3)] and the Elliott SU(3) model [\[5\]](#page-11-4), and they can be classified by their use of (a) quasiparticle single-particle states (b) mixing of particle-hole states and (c) schematic or realistic interactions. For example, the projected shell model (PSM) uses quasiparticle single-particle states but no mixing of particle-hole states and schematic interactions [\[6\]](#page-11-5); the MONSTER and VAMPIR codes use quasiparticle single-particle states, and mixing of particle-hole states, and realistic interactions [\[7\]](#page-11-6); and the projected configuration-interaction (PCI) eschews quasiparticles states but uses realistic interactions and mixes higher-order particle-hole states [\[8](#page-11-7)]. This list is not exhaustive.

What is new in this paper is the comparison of exact numerical solutions against a relatively simple state, with a focus on the quality of the excitation spectra. In this case the "exact" solution is from a full configuration-interaction (CI) calculation, using semi-realistic interactions diagonalized in a truncated but nontrivial shell-model basis, against a single Hartree-Fock state with angular momentum projection. Hence we eschew quasiparticle states and particle-hole states, but we use a realistic interaction and we allow arbitrary deformation.

This paper is similar in philosophy to previous work comparing the random phase approximation (RPA) against CI calculations in a CI basis $[9, 10]$ $[9, 10]$ $[9, 10]$, testing how good is a specific approximation against a numerically exact result; in the former work the approximation was RPA, while here it is PHF. Another similar set of papers are those on PCI [\[8\]](#page-11-7), which compare results directly to CI diagonalization. An important difference with much previous work, however, is that in addition to studying the trends of the ground state energies, we consider in detail the excitation spectra.

We find that rotational spectra, and spectra that are from simple particle-rotor coupling, are well reproduced, with a small rms error. Other excitation spectra, such as vibrational spectra or complex spectra from odd-A and odd-odd nuclides, have larger rms errors. This is unsurprising, but it is useful to verify in detail as we have done.

In the next two sections we outline the numerical exact (CI) and approximate (HF and PHF) calculations, which we then compare for a variety of nuclides in Section V

II. THE SHELL-MODEL BASIS AND CONFIGURATION-INTERACTION CALCULATIONS

We begin by outlining the shell-model basis in which we work and the fundamentals of configuration-interaction calculations [\[11](#page-11-10)[–13\]](#page-11-11).

We work entirely in occupation or configuration space. This means that we start with some finite set of orthonormal single-particle states, $\{\phi_a(\vec{r})\}$. Each of these states have good angular momentum and parity and are labeled by orbital l, total j, and z-component j_z; implicitly they have good spin s, which here is $1/2$ although other values are allowed by our code. We also have two species of particles, here protons and neutrons, each with fixed numbers.

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Using second quantization we have fermion creation and annihiliation operators $\hat{\phi}_a^{\dagger}, \hat{\phi}_a$ that create or remove a particle from the states $\phi_a(\vec{r})$. The uncoupled Hamiltonian is then

$$
\hat{H} = \sum_{ab} T_{ab} \hat{\phi}_a^{\dagger} \hat{\phi}_b + \frac{1}{4} \sum_{abcd} V_{ab,cd} \hat{\phi}_a^{\dagger} \hat{\phi}_b^{\dagger} \hat{\phi}_d \hat{\phi}_c \tag{1}
$$

The one-body and two-body matrix elements, T_{ab} and $V_{ab,cd}$ respectively, are appropriate integrals over the Hamiltonian and the single-particle states [\[11\]](#page-11-10). Because the Hamiltonian is an angular momentum scalar, we store the matrix elements in coupled form. Given the single-particle space, the matrix elements are generated externally to all our codes and read in from a file. Any radial form of the single-particle states and any form of the interaction, including non-local interactions, are allowed and make no pratical difference to our many-body codes. Because we work in second quantization, the two-body matrix elements are automatically antisymmeteric and we do not separate direct from exchange terms.

For configuration-interaction (CI) calculations we use the BIGSTICK code [\[14\]](#page-11-12). BIGSTICK creates a many-body basis of Slater determinants $\{|\alpha\rangle\}$, constructed from the single-particle basis:

$$
|\alpha\rangle = \prod_{a=1}^{N_p} \hat{\phi}_a^{\dagger} |0\rangle, \tag{2}
$$

where N_p is the number of particles (occupied states). We fix the total $M = J_z$ of the many-body basis states, (this is trivial as each single-particle state has good j_z as well) and is thus called an M-scheme code; because the Hamiltonian is an angular momentum scalar, the final eigenstates thus will automatically have good total angular momentum J (and isospin T, although we have the capability to break isospin; $T_z = (Z - N)/2$ is fixed). Within the fixed single-particle space and fixed J_z, T_z we allow all possible configurations.

Given the basis, BIGSTICK then computes the many-body Hamiltonian matrix elements $\langle \alpha | \hat{H} | \beta \rangle$ and finds the low-lying eigenstates, including the ground state, using the Lanczos algorithm [\[13,](#page-11-11) [15](#page-11-13)]. With the single-particle space and Hamiltonian matrix elements fixed, the resulting eigenenergies for the full-space CI calculation are numerically exact. BIGSTICK can handle CI M-scheme model spaces on a desktop computer up to dimension roughly 2-400 million.

III. HARTREE-FOCK APPROXIMATION

The Hartree-Fock approximation (HF) is a variational method using a single Slater determinant [\[1](#page-11-0)]. One applies a unitary transformation among the single-particle states:

$$
\hat{c}_i = \sum_{a=1}^{N_s} U_{ai} \hat{\phi}_a,\tag{3}
$$

where N_s is the number of single-particle basis states (hence the number of particles $N_p \lt N_s$) and creates the trial Slater determinant

$$
|\Psi_T\rangle = \prod_{i=1}^{N_p} \hat{c}_i^{\dagger} |0\rangle.
$$
 (4)

Then one finds a Slater determinant that minimizes the energy, that is, that minimizes

$$
E = \frac{\langle \Psi_T | \hat{H} | \Psi_T \rangle}{\langle \Psi_T | \Psi_T \rangle} \tag{5}
$$

Because our input matrix elements are antisymmeterized, we fully include both direct and exchange terms. Working in occupation space the exchange term causes us no difficulty (or, to put it another way, any difficulty is off-shored into calculation of the antisymmeterized integrals).

Many Hartree-Fock calculations enforce good angular momentum, for closed-shell systems or closed-shell plus or minus one particle. Our HF code (the SHERPA code [\[9\]](#page-11-8)) allows the Slater determinant to break rotational invariance, even for closed-shell systems; such Slater determinants are 'deformed.' Experience in nuclear physics suggest this to be a fruitful path. Our only constraint is that we assume the transformation [\(3\)](#page-1-0) to be real.

In order to minimize, we use the standard Hartree-Fock equations, which consist of iteratively solving

$$
\mathbf{h}\vec{u}_i = \epsilon_i \vec{u}_i,\tag{6}
$$

where the Hartree-Fock effective one-body Hamiltonian is

$$
h_{ab} = T_{ab} + \sum_{cd} V_{ac,bd} \rho_{dc} \tag{7}
$$

and

$$
\rho_{dc} = \sum_{i=1}^{N} U_{ic} U_{id} \tag{8}
$$

We have separate Slater determinants for proton and neutrons; the generalization is straight forward. More details are found in Appendix [A.](#page-6-0)

IV. PROJECTION OF ANGULAR MOMENTUM AND PARITY

In order to project out angular momentum, we introduce the standard projection operator $[1, 4]$ $[1, 4]$

$$
\hat{P}_{MM'}^J = \int d\Omega \mathcal{D}_{MM'}^{(J)*}(\Omega) \hat{R}(\Omega),\tag{9}
$$

where $\Omega = \alpha, \beta, \gamma$ and $d\Omega = d\alpha \sin\beta d\beta d\gamma$, $\mathcal{D}_{MM'}^{(J)*}$ is the Wigner D-matrix [\[16](#page-11-14)], and $\hat{R}(\Omega)$ is the rotation operator. The rotation operator acts on a Slater determinant U whose columns are single-particle states with good j, m; therefore the matrix elements of R are given by the Wigner D -matrix:

$$
\langle j'm'|\hat{R}|jm\rangle = \delta_{j'j}D_{m'm}^{j}(\alpha\beta\gamma).
$$
\n(10)

It is useful to note that the j, m, m' for the rotational matrix \hat{R} are those of the single-particle space, while the J, M, M' for the projection operator (9) are those of the many-body space.

We now introduce the Hamiltonian and "norm" matrix elements,

$$
h_{M,M'}^J \equiv \langle \Psi | \hat{H} \hat{P}_{M,M'}^J | \Psi \rangle, \ \ n_{M,M'}^J \equiv \langle \Psi | \hat{P}_{M,M'}^J | \Psi \rangle \tag{11}
$$

These are both Hermitian.

Then we solve the generalized eigenvalue equation

$$
\sum_{M'} h_{M,M'}^J g_{M'}^J = E^J \sum_{M'} n_{M,M'}^J g_{M'}^J.
$$
\n(12)

where we allow g_M to be complex. Although the deformed Hartree-Fock state will have an orientation, the final result will be independent of orientation; we confirmed this by arbitrarily rotating our HF state.

PHF will only generate a limited number of states. For a given J, the maximum number is $(2J + 1)$, although fewer can be found if the norm matrix n has zero, or very small, eigenvalues. This can happen, for example, if the HF state has axial symmetry. Therefore there will be states missing from the CI spectrum, such as low-lying excited $J=0$ states.

V. RESULTS

In this section we apply our calculations to a number of nuclides in the sd and pf shells as well as a limited p -sd space where we allow for parity-projection as well. We ask a number of questions:

• Because PHF is a variational theory, the PHF ground state energy will be above the exact ground state energy. We ask: what is the systematics of that displacement? For example, if the displacement were nearly constant, we could use PHF to reliably estimate the true ground state energy.

- In a similar fashion we ask, how well does PHF yield the excitation spectrum?
- For odd-A and odd-odd nuclei, how often does PHF yield the correct J^{π} assignment for the ground state?

• For multi-shell calculations, how well does parity+angular-momentum projection yield the splitting of positive and negative parity spectra?

Many previous studies have focused on the ground state energy. By contrast, we focus on the excitation spectrum, as well as systematics of the ground state energy.

We work with several model spaces and interactions. All model spaces assume some inert core and valence particles; the single-particle states can be thought of as spherical harmonic oscillator wavefunctions, but that has no impact on our calculations. The interactions we use are all 'realistic,' based on effective interactions derived from scattering data but with individual matrix elements adjusted to fit many-body data (see [\[11,](#page-11-10) [17](#page-11-15)] for details of methodology). The three model spaces we work in and their interactions are:

• sd, or the $1s_{1/2}$ -0d_{3/2}-0d_{5/2} valence space, assuming an inert ¹⁶O core; the interaction is the universal sd-interaction 'B,' or USDB [\[17\]](#page-11-15);

• pf, or the $1p_{1/2}$ - $1p_{3/2}$ - $0f_{5/2}$ - $0f_{7/2}$ valence space, assuming an inert ⁴⁰Ca core; the interaction is the monopolemodified Kuo-Brown G-matrix interaction version 3G, or KB3G [\[18](#page-11-16)];

• and finally the psd, or $0p_{3/2}$ - $0p_{1/2}$ - $0d_{5/2}$ - $1s_{1/2}$ valence space, assuming an inert ⁴He core; the interation is a hybrid of Cohen-Kurath (CK) matrix elements in the 0p shell[\[19\]](#page-11-17), the older universal sd interaction of Wildenthal [\[20](#page-11-18)] in the $0d_{5/2}$ -1s_{1/2} space, and the Millener-Kurath (MK) p-sd cross-shell matrix elements[\[21](#page-11-19)]. We leave out the $0d_{3/2}$ orbit is to make full shell-model calculations tractable. Within the p and sd spaces we use the original spacing of the single-particle energies for the CK and Wildenthal interactions, respectively, but then shift the sd single-particle energies up or down relative to the p-shell single particle energies to we get the first 3[−] state at approximately 6.1 MeV above the ground state. The rest of the spectrum, in particular the first excited 0^+ state, is not very good, but the idea is to have a non-trivial model, not exact reproduction of the spectrum. This model space and interaction allows us to consider model nuclei with both parities and to investigate parity-mixing in the HF state.

To illustrate our results, we begin with ²⁴Mg in the sd shell. Fig. [1](#page-4-0) compares the exact CI spectrum on the right against the PHF spectrum on the left. The PHF g.s. is 1.78 MeV above the CI g.s., while on average each states in the PHF spectrum is shifted by 1.68 MeV. For comparison, the correlation energy, here defined as the difference between the unprojected HF energy and the CI g.s. energy, is 6.62 MeV, so that restoring good J by projection accounts for 73% of the correlation energy.

Fig. [2](#page-4-1) is the same as Fig. [1](#page-4-0) but with the PHF spectrum shifted down so the g.s. energies coincide. The agreement between the excitation spectra is very good, with an rms error of 0.10 MeV.

Other nuclides shall qualitatively similar results, albeit with varying degrees of accuracy. In general the PHF excitation spectra of even-even nuclides was of much higher accuracy than for odd-odd and odd-A nuclides, which is not too surprising as one might expect a deformed HF state to approximate a rotational band.

For example, consider the odd-odd nuclide 30 Al in the sd shell, as shown in Fig. [3.](#page-5-0) The PHF g.s. is 6.40 MeV above the CI g.s., and PHF only accounts for 25% of the correlation energy. When the PHF g.s. is shifted down to the match the CI g.s., Fig. [4,](#page-5-1) the rms error in the excitation spectrum is 0.66 MeV and many of the states are out of order–in particular the first excited 1^+ state.

This continues in the pf shell. In Fig. [5](#page-6-1) we show 52 Ti, which has a strong vibrational rather than rotational spectrum. Here the PHF g.s. energy is 3.16 MeV above the CI g.s., account for only 42% of the correlation energy, and even when the PHF spectrum is shifted down in Fig. [6,](#page-6-2) it is clearly compressed relative to the exact CI, with an rms error of 0.53 MeV.

By contrast, the spectrum of ⁴⁹Cr, which can be thought of as a particle-rotor coupled nucleus, is good. As shown in Fig. [7,](#page-7-0) the PHF g.s. is 2.63 MeV above the CI g.s., and PHF accounts for only 35% of the correlation energy; but when the PHF spectrum is shifted down, as in Fig. [8,](#page-7-1) the rms error is only 0.01 MeV.

We also worked in a cross-shell system, the $p-gd_{5/2}$ space. Keep in mind we artificially reduced the separation between the p and $sd_{5/2}$ shells in order to provoke a parity-mixed HF state.

The vibrational spectrum the positive parity states of ¹⁸C is only modestly reproduced by PHF, as illustrated in Fig. [9](#page-8-0) and [10.](#page-8-1) The PHF g.s. is 3.84 MeV above the CI g.s., accounting for only 44% of the correlation energy, and when the PHF spectrum is shifted to match the g.s. energies, the rms error is 0.22 MeV , with the 4^+ state in the wrong place.

Somewhat better are the positive parity states of ^{22}F , in Figs. [11](#page-9-0) and [12.](#page-9-1) The PHF state is 1.41 MeV above the CI g.s., but that only accounts for barely 15% of the correlation energy. On the other hand, when the g.s. energies are matched, the error in the excitation is only 0.08 MeV, although some of the states are in the wrong order.

Tables I and II summarize our results. In particular, Table I summarizes for specific cases the standard deviation (rms error) of the PHF excitation spectra relative to the CI excitation spectra, once the ground state energies are matched. Table II summarizes all our calculations; the rightmost column gives how often the correct g.s. J^{π} arises out of the PHF. Even-even nuclides almost always get the correct $J = 0$ ground state (but for our mixed-shell calculations, occasionally the wrong parity–but keep in mind we artificially reduced the splitting between the shells

FIG. 1: Comparison of exact CI and PHF spectra of ^{24}Mg in the sd model space with the USDB interaction.

FIG. 2: Comparison of exact CI and PHF spectra of ²⁴Mg in the sd model space with the USDB interaction, but with the PHF spectrum shifted downwards so the g.s. energies coincide.

to force a mixed-parity HF state), but odd-A got the correct g.s. assignment only a little over half the time, and odd-odd nuclei less than half the time.

VI. CONCLUSIONS

We have carried out detailed angular momentum projected Hartree-Fock calculations in a shell-model basis with semi-realistic shell-model interactions, and compared the spectra to exact configuration-interaction diagonalization calculations, with a particular focus on the quality of the excitation spectra. We found, unsurprisingly, that rotational nuclei are best reproduced by PHF.

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FIG. 3: Comparison of exact CI and PHF spectra of ³⁰Al in the sd model space with the USDB interaction.

FIG. 4: Comparison of exact CI and PHF spectra of ³⁰Al in the sd model space with the USDB interaction, but with the PHF spectrum shifted downwards so the g.s. energies coincide.

				Valence Std. dev. % PHF	
Shell	Nucleus ^[Z]		N	(MeV)	states
$_{sd}$	24 Mg	4	$\overline{4}$	0.077	
	30 Al	$\overline{5}$	9	1.004	12
p f	52T _i	$\overline{2}$	10	0.714	8
	^{49}Cr	4	5	0.074	4
$p\hbox{-} sd_{5/2}$	$^{18}\mathrm{C}$ $^{22}\mathrm{F}$	4	10	0.810	
		7	11	0.194	5

TABLE I: Summary of data for the nuclei in the paper. The standard deviation is for the shifted PHF spectra relative to the exact CI spectra.

FIG. 5: Comparison of exact CI and PHF spectra of 52 Ti in the pf model space with the KB3G interaction.

FIG. 6: Comparison of exact CI and PHF spectra of 52 Ti in the pf model space with the KB3G interaction, but with the PHF spectrum shifted downwards so the g.s. energies coincide.

Appendix A: Some computational details for Hartree-Fock

A single general Slater determinant we write as a rectangular, $N_s \times N_p$ matrix Ψ . Operationally, the vectors \vec{u}_i , of length N_s , from solving the Hartree-Fock equation [\(6\)](#page-2-1) form the columns of the transformation of the single-particle basis [\(3\)](#page-1-0). We take the N_p vectors with lowest ϵ_{λ} to form the columns of Ψ . In other words, Ψ is simply selecting columns of U, which are computed in Eq. (6) ; this is the unitary transformation that diagonalizes **h**. We have separate proton and neutron Slater determinants, Ψ_{π} and Ψ_{ν} .

From the Slater determinant we construct the one-body density matrix,

$$
\rho_{ab} = \left\langle \Psi \left| \hat{\phi}_a^{\dagger} \hat{\phi}_b \right| \Psi \right\rangle = \sum_{i=1}^{N_p} \Psi_{ai} \Psi_{bi}^* . \tag{A1}
$$

We then construct the Hartree-Fock effective one-body Hamiltonian,

$$
h_{ab} = H_{ab}^1 + \sum_{cd} V_{ac,bd} \rho_{cd}.
$$
 (A2)

FIG. 7: Comparison of exact CI and PHF spectra of ⁴⁹Cr in the pf model space with the KB3G interaction.

FIG. 8: Comparison of exact CI and PHF spectra of ⁴⁹Cr in the pf model space with the KB3G interaction, but with the PHF spectrum shifted downwards so the g.s. energies coincide.

Because we have two species, we have to take a little care. If our interaction is an isospin scalar, then

$$
V_{abcd}^{pp} = V_{abcd}^{nn} = V_{abcd}^{T=1}
$$
\n(A3)

and

$$
V_{abcd}^{pn} = \frac{1}{2} \left(V_{abcd}^{T=0} + V_{abcd}^{T=1} \right). \tag{A4}
$$

Here we have suppressed the obvious decoupling of angular momentum [\[11,](#page-11-10) [16\]](#page-11-14).

Then the proton Hartree-Fock Hamiltonian is

$$
h_{ab}^{\pi} = H_{ab}^{1} + \sum_{cd} V_{ac,bd}^{pp} \rho_{cd}^{\pi} + V_{ac,bd}^{pn} \rho_{cd}^{\nu}
$$
 (A5)

and similarly for the neutron Hartree-Fock Hamiltonian.

FIG. 9: Comparison of exact CI and PHF spectra of ¹⁸C in the p -sd_{5/2} model space.

FIG. 10: Comparison of exact CI and PHF spectra of ¹⁸C in the p -sd_{5/2} model space, but with the PHF spectrum shifted downwards so the g.s. energies coincide.

Appendix B: Some computational details for projected Hartree-Fock

We now have to compute $\langle \Psi | \hat{R}(\Omega) | \psi \rangle$ and $\langle \Psi | \hat{H} \hat{R}(\Omega) | \psi \rangle$, where \hat{R} is the rotation operator. In our case

$$
\hat{R}|\Psi\rangle \to \mathbf{R}\Psi = \tilde{\Psi} \tag{B1}
$$

where Ψ and $\tilde{\Psi}$ are $N_s \times N_p$ matrices representing our original and rotated Slater determinants, respectively, and R is the square matrix the carries out the rotation. In our single-particle basis the matrix elements are straightforward:

$$
R_{ab} = \langle j_a m_a | \hat{R}(\alpha, \beta, \gamma) | j_b m_m \rangle = \delta_{j_a j_b} \mathcal{D}^{j_a}_{m_a, m_b}(\alpha, \beta, \gamma), \tag{B2}
$$

where $\mathcal{D}_{m',m}^j$ is the Wigner D-matrix [\[16\]](#page-11-14).

Compute the matrix elements between two oblique Slater determinants is relatively straightforward [\[22\]](#page-11-20). First,

$$
\langle \Psi | \hat{R} | \Psi \rangle = \langle \Psi | \tilde{\Psi} \rangle = \det \Psi^{\dagger} \tilde{\Psi}.
$$
 (B3)

FIG. 11: Comparison of exact CI and PHF spectra of ²²F in the p -s $d_{5/2}$ model space.

FIG. 12: Comparison of exact CI and PHF spectra of ²²F in the $p-sd_{5/2}$ model space, but with the PHF spectrum shifted downwards so the g.s. energies coincide.

Calculation of the Hamiltonian (and other) expectation value requires the density matrix,

$$
\rho_{ab}' = \frac{\langle \Psi | \hat{\phi}_a^\dagger \hat{\phi}_b | \tilde{\Psi} \rangle}{\langle \Psi | \tilde{\Psi} \rangle} = \left[\tilde{\Psi} \left(\Psi^\dagger \tilde{\Psi} \right)^{-1} \Psi^\dagger \right]_{ba} \tag{B4}
$$

Then

$$
\langle \Psi | \hat{H} | \tilde{\Psi} \rangle = \sum_{ab} H_{ab}^1 \rho'_{ab} + \sum_{abcd} V_{ab,cd} (\rho'_{ac} \rho'_{bd} - \rho'_{ad} \rho'_{bc})
$$
(B5)

Accounting for two species is straightforward. Note that even if our original Slater determinants were real, by rotating over all angle they can become complex.

We only project out good angular momentum. We could in principle project out exact isospin; to do this we would have to use a single Slater determinant containing both protons and neutrons and rotate in isospin space. We leave such a modification for future work.

					No. Correct Incorrect Frequency
$Z - N$			Shell nuclei G.S. J	G.S. J	correct
Even-Even	$s-d$	15	15	0	1.00
	$p-f$	10	10	θ	1.00
	p -sd	17	12	5	0.71
	$s-d$	15	8	7	0.53
Odd-Odd	$p-f$	6	$\overline{2}$	4	0.33
	p -sd	12	4	8	0.33
	$s-d$	25	17	8	0.68
Odd-A	$p-f$	11	6	5	0.55
	p -sd	25	13	12	0.52

TABLE II: Frequency of ground state matching. The "No. nuclei" column represents the total number of nuclei that was studied for that shell that satisfies the Z-N relationship.

Appendix C: Linear algebra of projected states

In order to calculate the weights of the angular momentum-projected states in the Hartree-Fock state, we have to investigate in some detail the linear algebra of solving Eq [\(12\)](#page-2-2).

Specifically, we want to expand the Hartree-Fock state

$$
|HF\rangle = \sum_{\lambda} a_{\lambda} |\lambda\rangle.
$$
 (C1)

Now each state λ has definite J, but can be broken up into M-components. Thus we expand

$$
|HF\rangle = \sum_{\lambda,M} a_{\lambda,M} |\lambda,M\rangle \tag{C2}
$$

and we ultimately want

$$
|a_{\lambda}|^2 = \sum_{M} |a_{\lambda,M}|^2. \tag{C3}
$$

As discussed above, we want to solve the generalized eigenvalue problem

$$
\mathcal{H}\vec{g}_{\lambda} = E_{\lambda}\mathcal{N}\vec{g}_{\lambda} \tag{C4}
$$

Now because

$$
\sum J K M \hat{P}_{K,M}^{J} = 1
$$
\n(C5)

we have $tr \mathcal{N} = 1$.

To solve the generalized eigenvalue problem we decompose the norm matrix, $\mathcal{N} = S S^{\dagger}$, for example by Cholesky or by diagonalization. In general there will be a null space which we project out.

By creating $h = S^{-1} \mathcal{H} S^{-1\dagger}$ (and also projecting out the null space, to eliminate division by zero), we then solve the ordinary eigenvalue problem

$$
\mathbf{h}\vec{v}_{\lambda} = E_{\lambda}\vec{v}_{\lambda}.\tag{C6}
$$

Of course, $|\vec{v}_\lambda|^2 = 1$. Furthermore, because of this, the vectors $\vec{g}_\lambda = mathbf{S}^{-1\dagger} \vec{v}_\lambda$ are orthonormal with respect to the norm matrix, that is

$$
\vec{g}_{\lambda}^{\dagger} \mathcal{N} \vec{g}_{\mu} = \delta_{\lambda \mu}.
$$
 (C7)

Another way to see this is to note that

$$
|\lambda, JM\rangle = \sum_{K} \hat{P}_{M,K}^{J} |HF\rangle g_K(\lambda)
$$
 (C8)

and because $\hat{P}^2 = \hat{P}$ we automatically get [\(C7\)](#page-10-0).

As stated above, our ultimate goal is to compute $|a_{\lambda}|^2$. Towards this end, we note that

$$
a_{\lambda,M} = \langle \lambda, JM | HF \rangle = \sum_{K} \mathcal{N}_{M,K}^{J} g_K(\lambda)
$$
 (C9)

so that

$$
|a_{\lambda}|^2 = \sum_{M} |a_{\lambda,M}|^2 = \sum_{K,K'} g_K^*(\lambda) (\mathcal{N}^2)_{K,K'}^2 g_{K'}(\lambda) = \vec{g}_{\lambda}^{\dagger} \mathcal{N}^2 \vec{g}_{\lambda}.
$$
 (C10)

We can rewrite this as

$$
|a_{\lambda}|^2 = \vec{v}_{\lambda}^{\dagger} \mathcal{S}^{\dagger} \mathcal{S} \vec{v}_{\lambda}
$$
 (C11)

which is what we want in convenient terms. To confirm this is correct, we compute

$$
\sum_{\lambda} |a_{\lambda}|^2 = \sum_{\lambda} \vec{v}_{\lambda}^{\dagger} \mathcal{S}^{\dagger} \mathcal{S} \vec{v}_{\lambda} = \text{tr}\left(\mathcal{S} \sum_{\lambda} \vec{v}_{\lambda} \vec{v}_{\lambda}^{\dagger} \mathcal{S}^{\dagger}\right)
$$
(C12)

but using the completeness relation, the sum over λ yields just 1, and we have

$$
\text{tr} \mathcal{S} \mathcal{S}^{\dagger} = \text{tr} \, \mathcal{N} = 1 \tag{C13}
$$

as we'd expect.

- [1] P. Ring and P. Shuck, The nuclear many-body problem (Springer-Verlag, New York 1980).
- [2] R. E. Peierls and J. Yoccuz, Proc. Phys. Soc. (London) A70, 381 (1957).
- [3] S. G. Nilsson, Mat. Fys. Medd. Dan. Vid. Selsk.29, No. 16 (1955); A. Bohr and B. R. Mottelson, Nuclear structure, volume II: nuclear deformations (W. A. Benjamin, New York, 1975).
- [4] R. D. Lawson, Theory of the nuclear shell model (Clarendon Press, Oxford, 1980).
- [5] J. P. Elliot, Proc. Roy. Soc. (London) A245, 128 (1958); J. P. Elliot, Proc. Roy. Soc. (London) A245, 562 (1958); I. Talmi, Simple models of complex nuclei (Harwood Academic Publishers, Chur, Switzerland, 1993).
- [6] K. Hara and Y. Sun, Int. J. Mod. Phys. E 4, 637 (1995); J.A. Sheikh and K. Hara, Phys. Rev. Lett. 82, 3968 (1999); Y.S. Chen and Z.C. Gao, Phys. Rev. C 63, 014314 (2001); J.A. Sheikh, Y. Sun and R. Palit, Phys. Lett. B 507, 115 (2001); Y. Sun, J.A. Sheikh and G.-L. Long, Phys. Lett. B 533, 253 (2002); Y. Sun and C.-L. Wu, Phys. Rev. C 68, 024315 (2003), p. 024315.
- [7] K. W. Schmid, F. Grümmer, and A. Faessler, Phys. Rev. C 29, 291 (1984); E. Bender, K. W. Schmid, and A. Faessler, Phys. Rev. C 52, 3002 (1995); K. W. Schmid, Prog. Part. Nucl. Phys. 52, 565 (2004).
- [8] Z.-C. Gao and M. Horoi, Phys. Rev. C 79, 014311 (2009); Z.-C. Gao, M. Horoi, and Y. S. Chen, Phys. Rev. C 80, 034325 (2009)
- [9] I. Stetcu, PhD. thesis, Louisiana State University, 2003,<http://etd.lsu.edu/docs/available/etd-0702103-154150/>
- [10] I. Stetcu and C.W. Johnson, Phys. Rev. C 66, 034301 (2002).
- [11] P.J. Brussard and P.W.M. Glaudemans, Shell-model applications in nuclear spectroscopy (North-Holland Publishing Company, Amsterdam 1977).
- [12] B. A. Brown and B. H. Wildenthal, Annu. Rev. Nucl. Part. Sci. 38, 29 (1988).
- [13] E. Caurier, G. Martinez-Pinedo, F. Nowacki, A. Poves, and A. P. Zuker, Rev. Mod. Phys. 77, 427 (2005).
- [14] C. W. Johnson, W. E. Ormand, and P. G. Krastev, [arXiv:1303.0905.](http://arxiv.org/abs/1303.0905)
- [15] R. R. Whitehead, A. Watt, B. J. Cole, and I. Morrison, Adv. Nucl. Phys. 9, 123 (1977).
- [16] A. R. Edmonds, Angular Momentum in Quantum Mechanics, (Princeton University Press, Princeton, 1960).
- [17] B.A. Brown and W.A. Richter, Phys. Rev. C 74 034315 (2006).
- [18] A. Poves, J. Sánchez-Solano, E. Caurier and F. Nowacki, Nucl. Phys. A 694, 157 2001.
- [19] S. Cohen and D. Kurath, Nucl. Phys. **73**, 1 (1965).
- [20] B.H. Wildenthal, Prog. Part. Nucl. Phys. 11, 5 (1984).
- [21] D. J. Millener and D. Kurath, Nucl. Phys **A 255**, 315 (1975).
- [22] G. H. Lang, C. W. Johnson, S. E. Koonin, and W. E. Ormand, Phys. Rev. C 48, 1518 (1993).